



Greyline Propagation of LF/HF signals

***- Fact and Fiction (or do
observations fit the theory?)***

 Steve Nichols G0KYA, RSGB Propagation Studies Committee

This was the title of a talk given to the RSGB HF Convention on Saturday 12th October 2002. For further, basic, information please refer to the article on Greyline Propagation written by Steve Nichols G0KYA and published in the RSGB's June 2002 issue of RadCom



What is “Greyline” propagation?

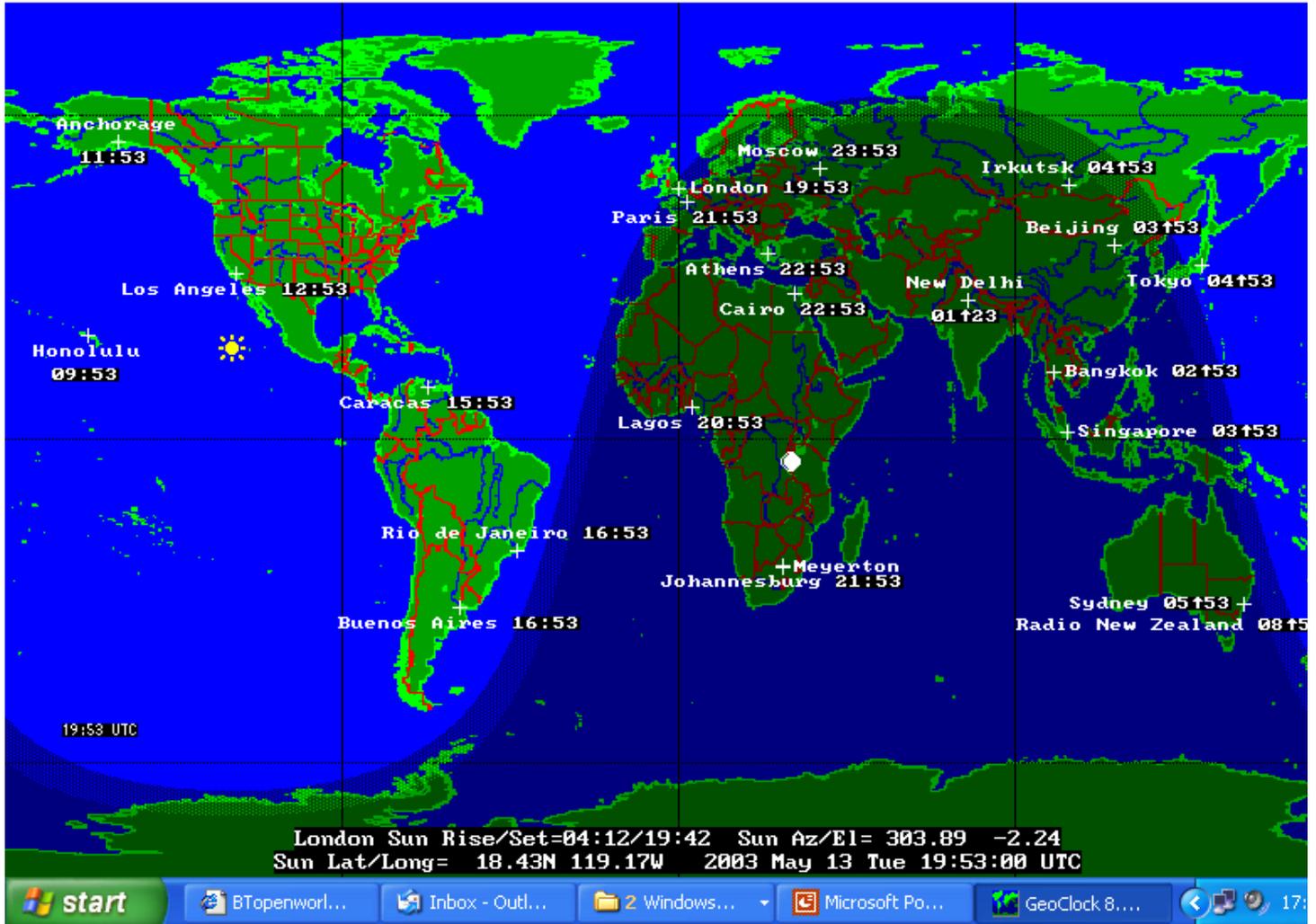
- The phrase was first coined in 1975
- It describes the propagation of radio waves **along** the terminator separating night and day (sunrise/sunset)
- this is the “Greyline” path (or “Grayline” in USA)
- The path is reported to be extremely efficient
- It is not always reported as an LF effect

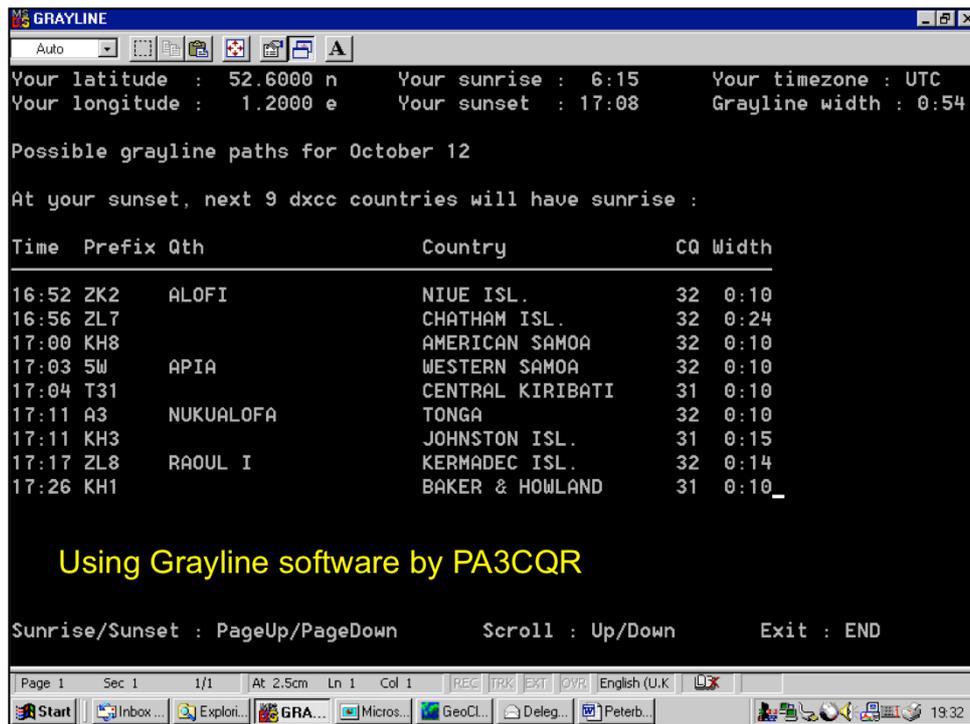
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The main point was that all books on propagation refer to greyline without being too specific about the bands actually affected.

One article, in Practical Wireless discussed greyline propagation 28MHz. As greyline is basically described as being the zone where D layer absorption is at a minimum, while F layer illumination is still very high, it is unlikely that 28Mhz would see any benefit.

However, 28MHz paths crossing the terminator may experience skip focusing effects - not to be confused with greyline which is along the terminator.





This software by PA3CQR can calculate supposed greyline openings by calculating simultaneous sunrise/sunset.



The D Layer:

- Is at an average height of about 60-90km
- It is mainly responsible for absorbing/attenuating signals on 160-20m. It can reflect signals at VLF
- This is why we don't hear "much" skip on the LF bands during daylight, especially in summer
- The actual level of absorption is dependent upon the frequency, time of day/year and solar/geomagnetic levels





D Layer Absorption:

$$\text{Absorption (db)} = (10 \cdot \log[\text{flux (W m}^{-2}\text{)}] + 65) \cdot 2/f^2$$



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The main point is that absorption goes as the inverse square of the frequency



Relative D Layer Absorption:

According to the RSGB "*Amateur Radio Operating Manual*", for a daytime path:

- 6db attenuation per hop on 21MHz
- 40db attenuation per hop on 7 MHz
- 100db attenuation per hop on 3.5MHz



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Thoughts from KN4LF (160m)

- The D layer does not always disappear at night
- It is affected by flares, CMEs, Sporadic-D, solar background X-ray level and energetic proton events
- The angle of the signal can have a profound effect (Magneto Ionic Power Coupling)
- Polar cap absorption can be severe for signals going near the poles. Also, consider electron gyro frequency absorption effects.

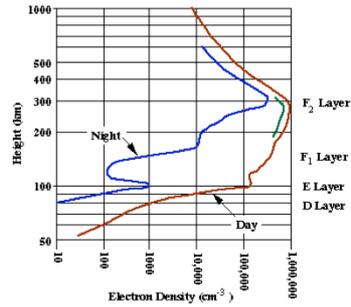


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So why is the Greyline thought to be efficient?

- At sunset, the lower D layer is no longer illuminated, therefore absorption is reducing
- The higher F-layer(s) are still illuminated, so we still have ionisation to aid propagation
- At sunrise, the reverse occurs





So what can we hear at sunrise?

Theoretically,

- ZS6 (South Africa) in August
- ZL (North Island, New Zealand) in December
- LU/PY/VP8 (Argentina, Brazil, Falklands) in December
- And others according to the season



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These are theoretical



So what can we hear at sunset?

Theoretically,

- JA (Japan) in August
- ZL (New Zealand) in September
- ZS6 (South Africa) in November
- LU/PY/VP8 (Argentina, Brazil, Falklands) in June
- And others according to the season



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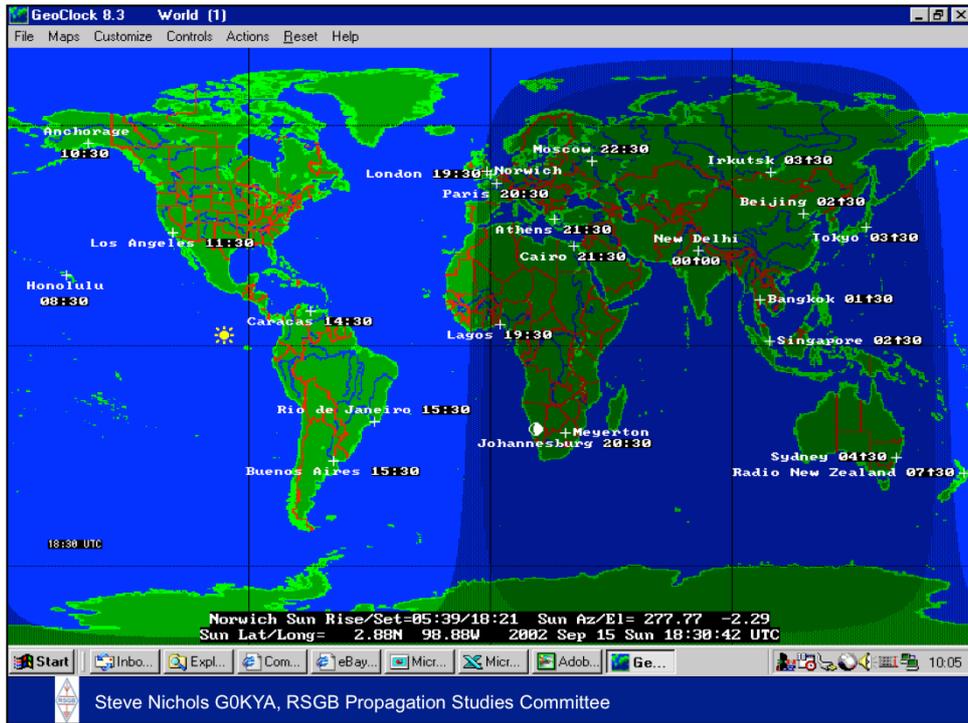
These are theoretical

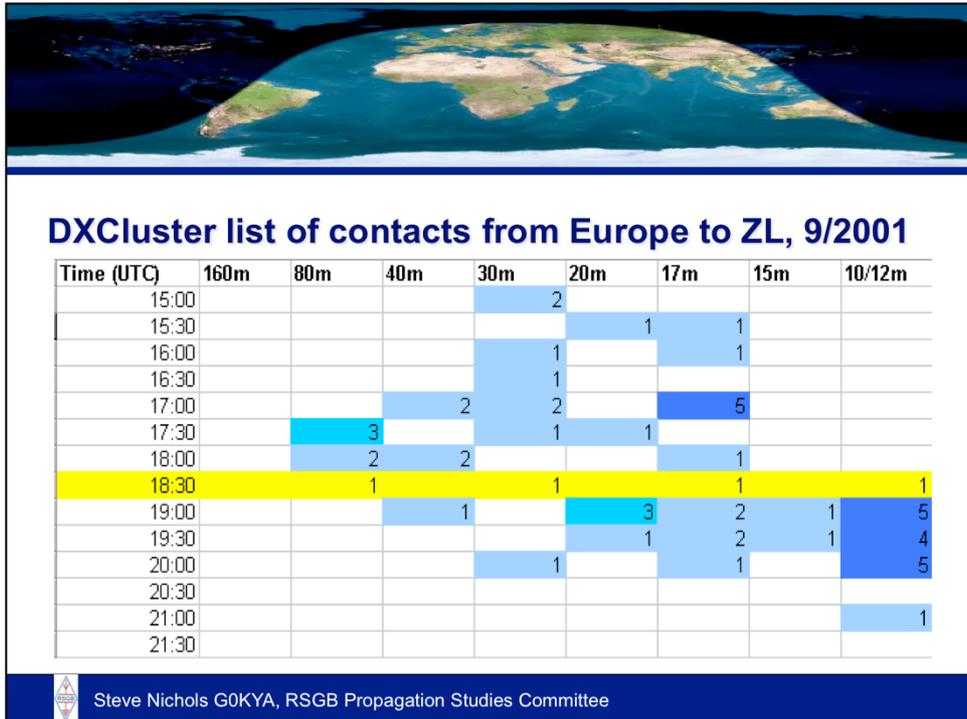


So let's take a look!



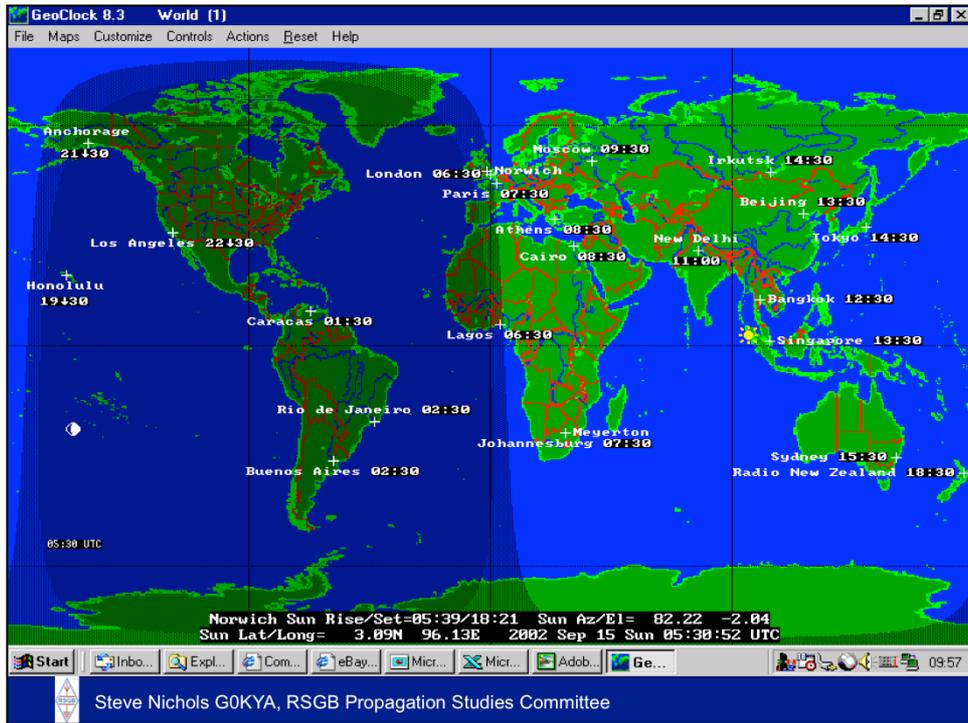
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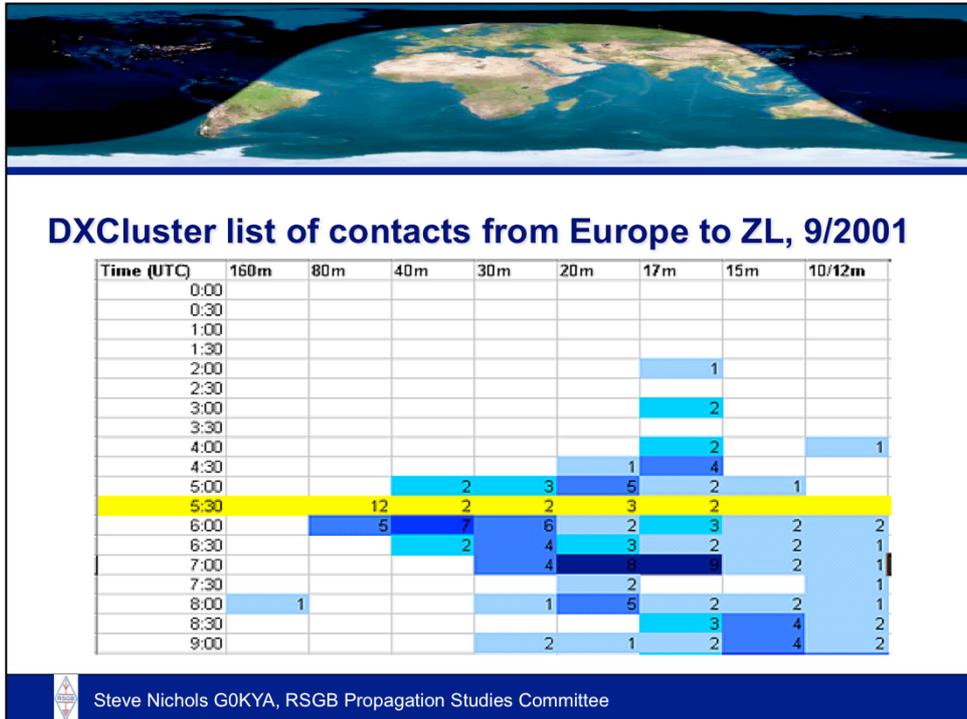




This was an analysis of the DX cluster of all contacts into ZL from UK and surrounding countries for September 2001.

The yellow line shows the greyline - as you can see there were few contacts, showing that greyline openings are not as common as we are led to believe.





There are however many contacts at the sunset condition on 80m. The previous map does not however show a true greyline path.



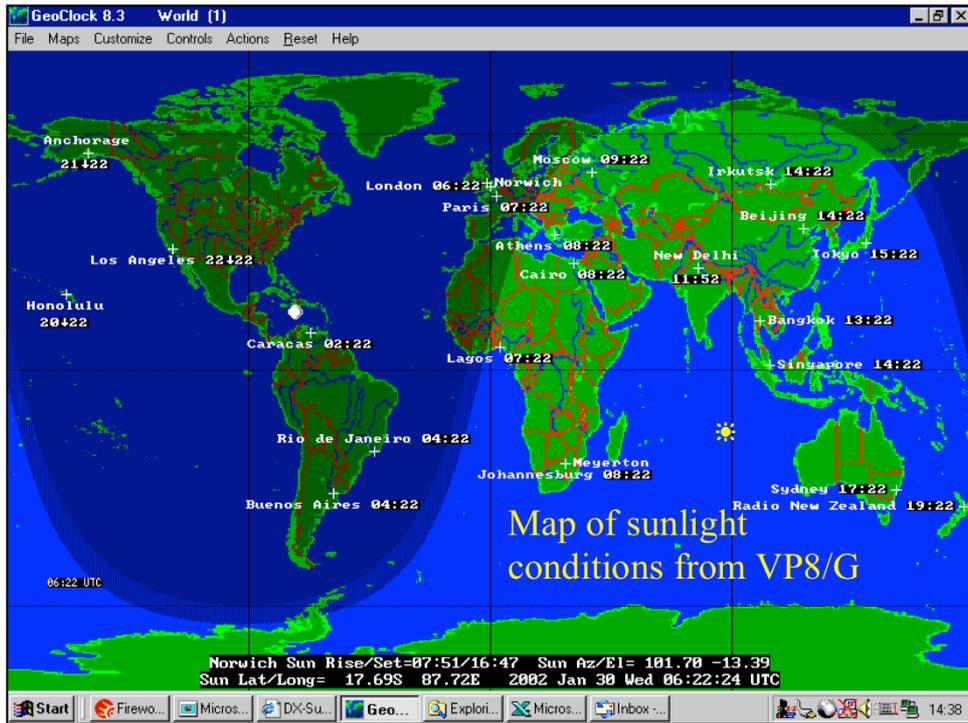
So what about about G/VP8?

- On 40m/80m there were no logged contacts on the DXCluster in December and June 2001 at all
- All contacts that did take place were definitely outside the Greyline times
- But this one contact stands out:

G3MLO 7077.0 VP8GEO **PEAKING S9 IN UK**
06:22 30 Jan 2002



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The contact into VP8 South Sandwich Island was not a true greyline path using sunrise/sunset as our measure of the greyline. My argument is that what we are interested in is what is happening above our heads at 75km - not what is happening on the ground.

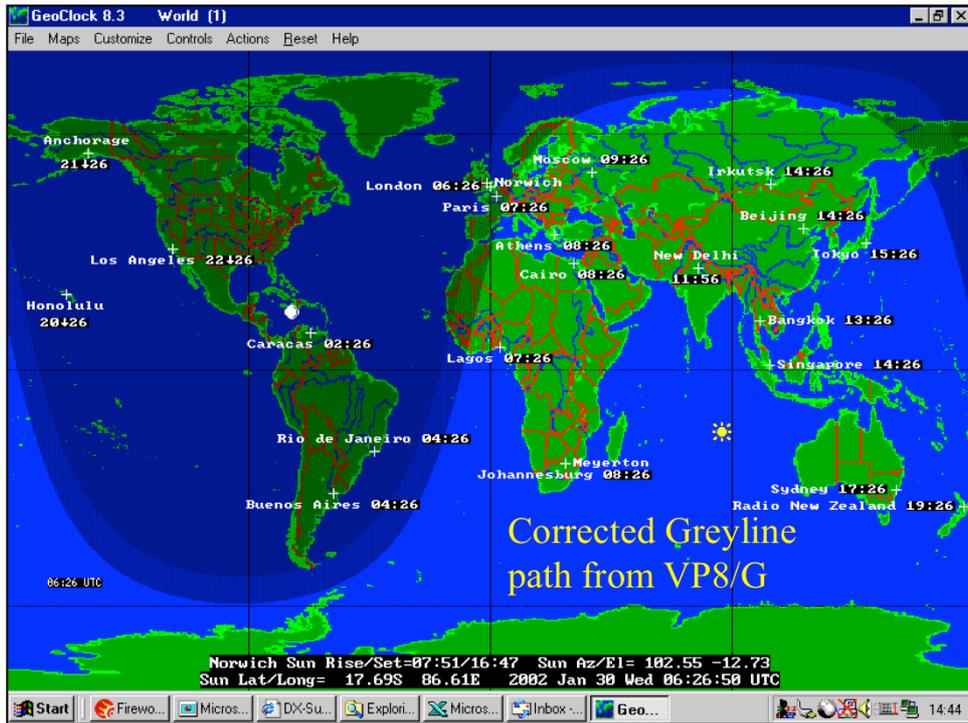


Is the Greyline where we think it is?

- The D layer is at a height of about 60-90km
- The F layer(s) are at a height of 150-300km
- So, sunrise/sunset tells us nothing about the illumination of the D/F layers, which is what we need to know
- By changing Geoclock to show the start/stop of illumination of the D/F layers we get



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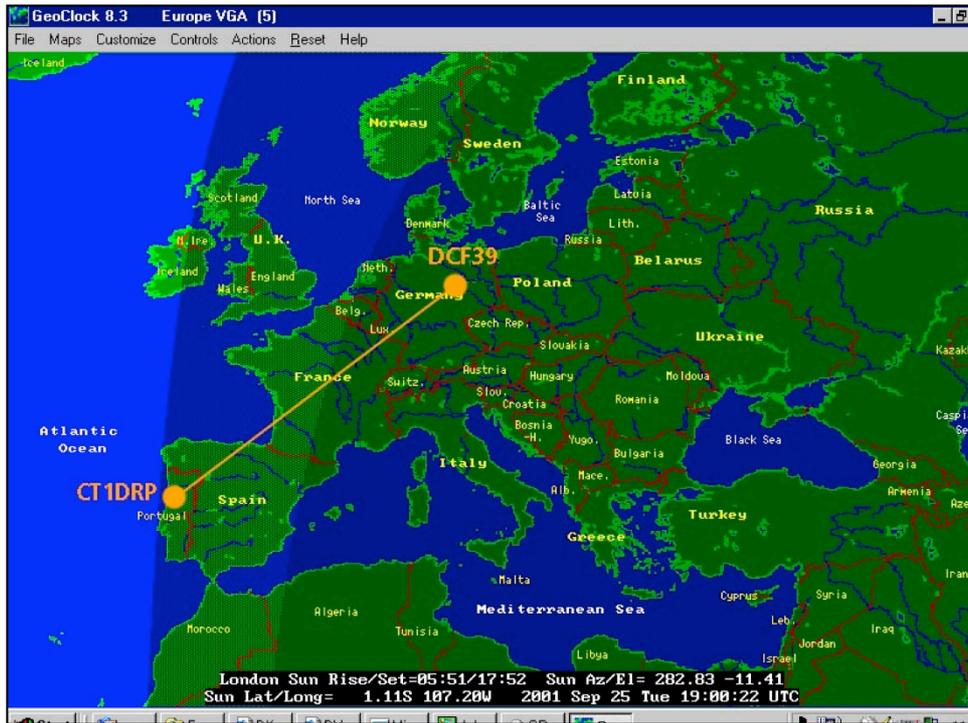
By showing the D/F layer illumination we get a better picture of where the greyline is.



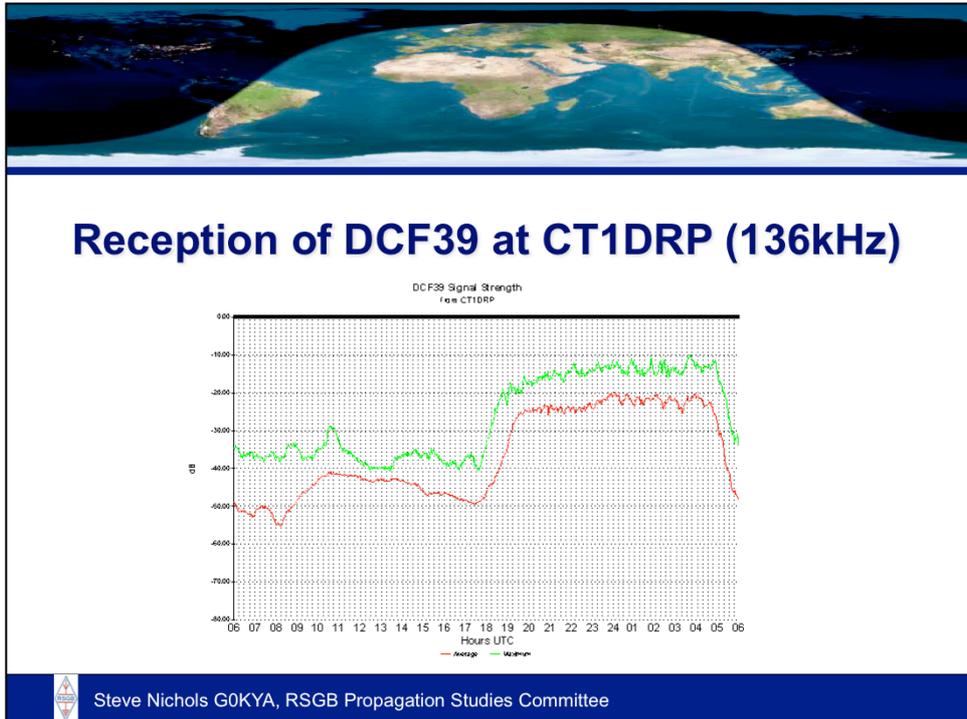
**Is this model of the
D layer illumination correct?**



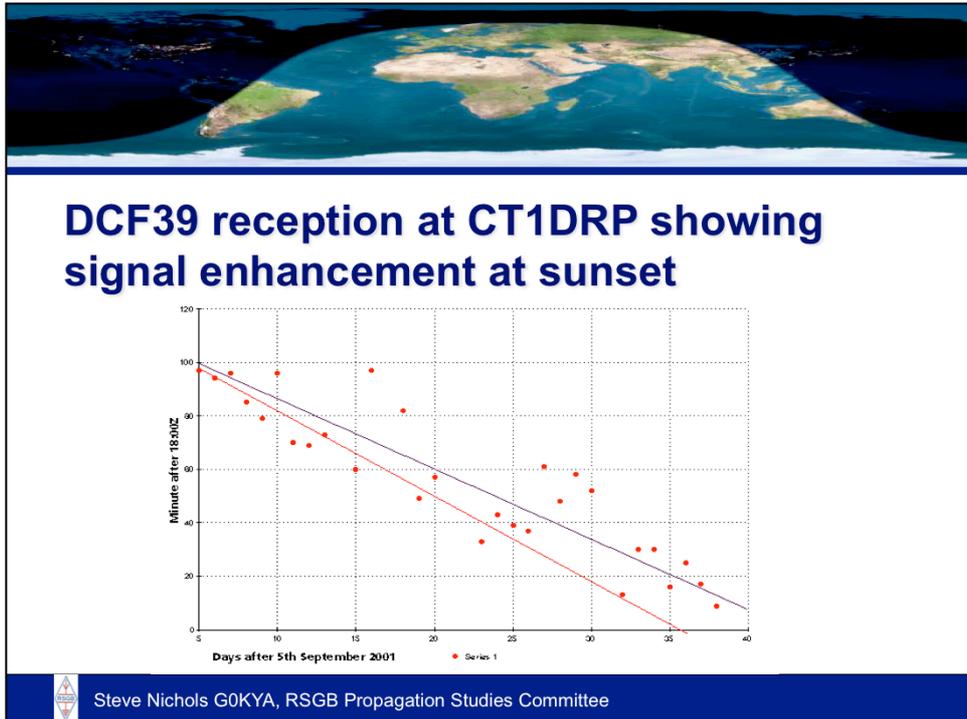
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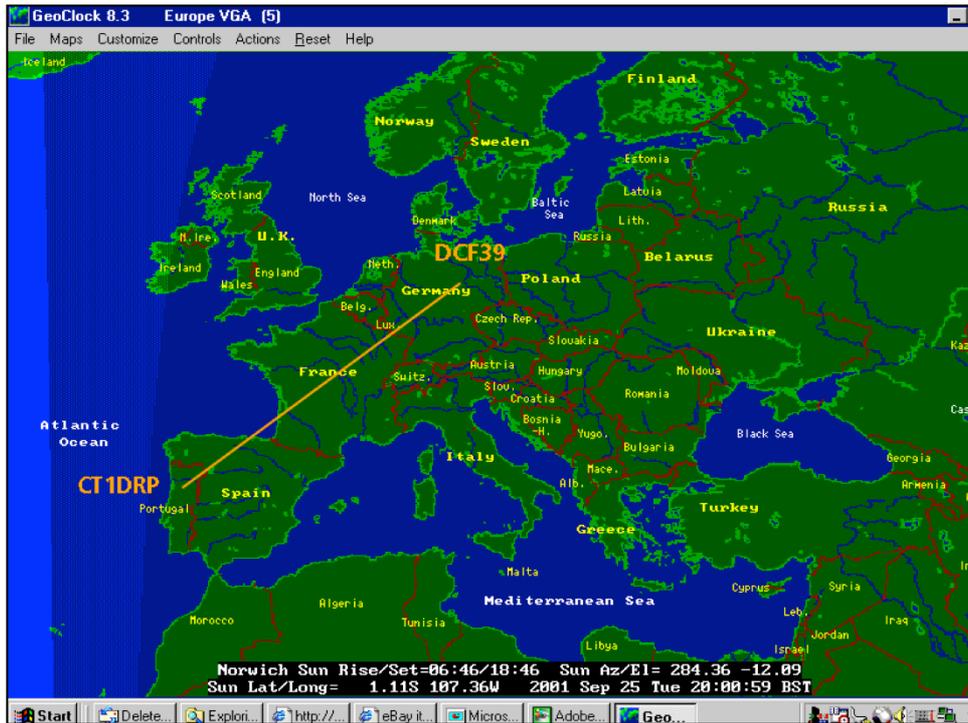
We have extensive data for the reception of the DCF39 into Portugal. The above shows the uncorrected sunset map for the path.



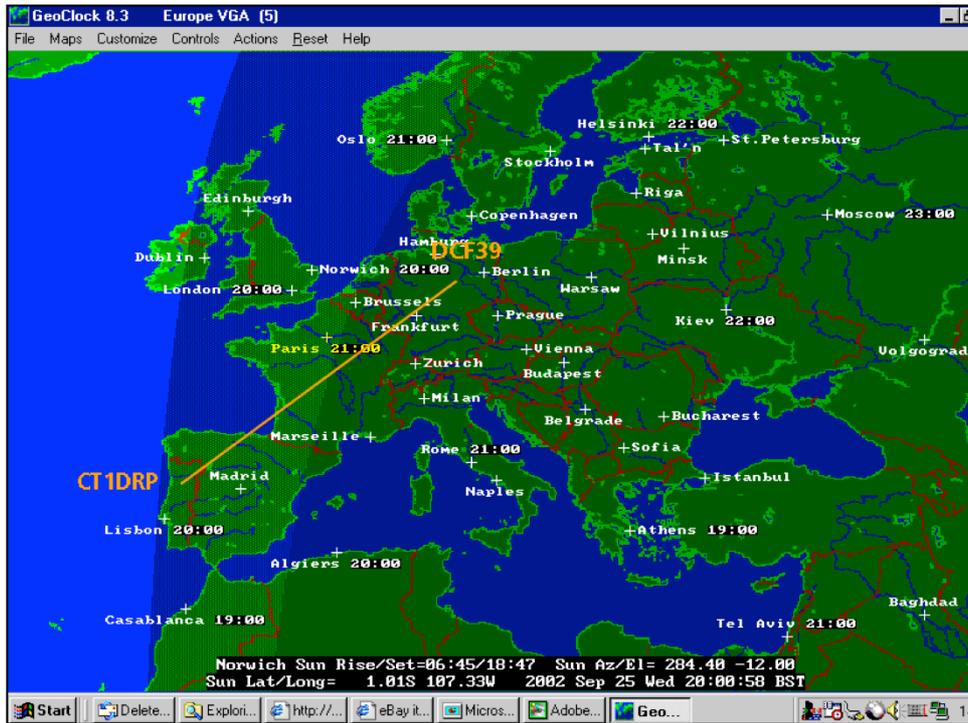
The upturn is due to the D layer illumination ending some time after sunset. Before that time the signal bounces off the bottom of the D layer. After that time is bounce off the bottom of the E layer. We can use this data to map where the D layer absorption ends.



This graph shows that this is not that predictable - it varies from day to day even allowing for sunset occurring earlier. In other words, the data points don't lay on a straight line.



This shows the sunset condition - it doesn't explain the upturn in signal. There is little correlation.



This shows the adjusted greyline. Now we can see that the signal increases some time after the D layer illumination ends and just before F layer illumination ends. We can now test this hypothesis by looking for greyline enhancement and checking any upturn with the time.



But the “corrected” path does not appear to work with all so-called greyline contacts, especially those from VK, ZL etc

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This theory seems to hold for local paths, but not for long paths. More work needed!



The purpose of Project Greyline

- To find better ways of modelling the greyline
- To find out if greyline paths actually exist in the way we currently model them
- To find out what frequencies are the most efficient over greyline paths if any.
- To look at trans-terminator paths (skip focusing)

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This is the new group that has been set up. The home page can be found at <http://uk.groups.yahoo.com/group/greyline/>



The experiments currently underway

- Meyerton - reception of the BBC World Service from South Africa on 3255kHz ✕
- Globe Wireless SITOR experiments ✕
- Chirp Sounder on 5MHz - looking for RTW ✕
- VP8 – G on 40m 👍
- ZL – G on 80m 👍

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We are looking at the reception of signals from South Africa in early November to look for any greyline enhancement. Since the lecture a station in Lusaka has been added too on 4265KHz.

We also have a list of frequencies and location for a worldwide network of SITOR stations

We are also looking for round the world greyline echoes on 5MHz using chirpsounder - none found so far.

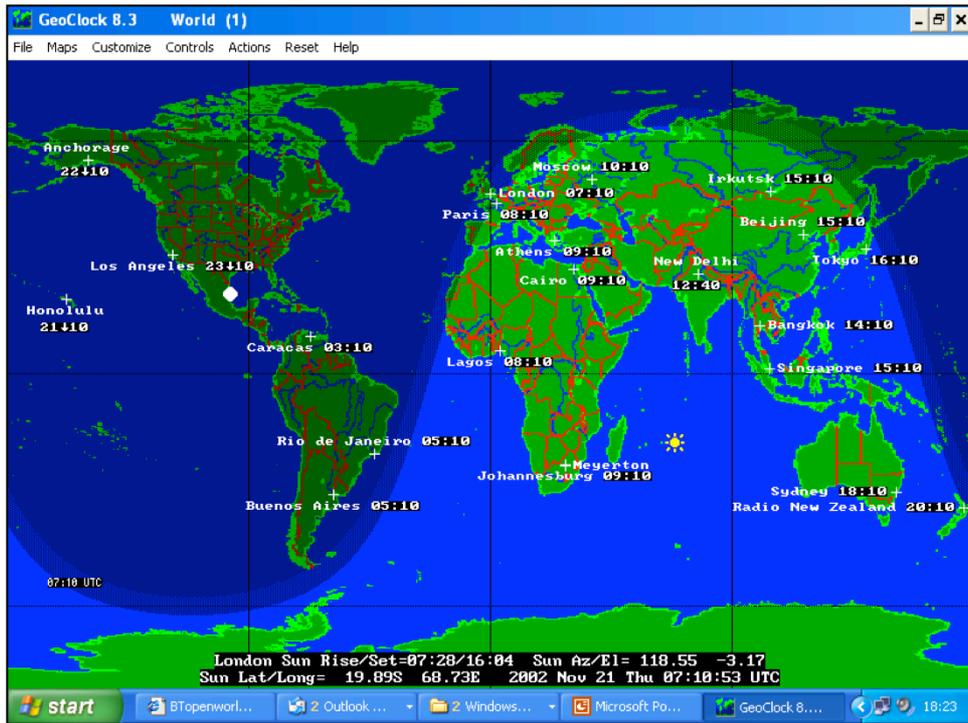


The G<>VP8 experiment 11/02

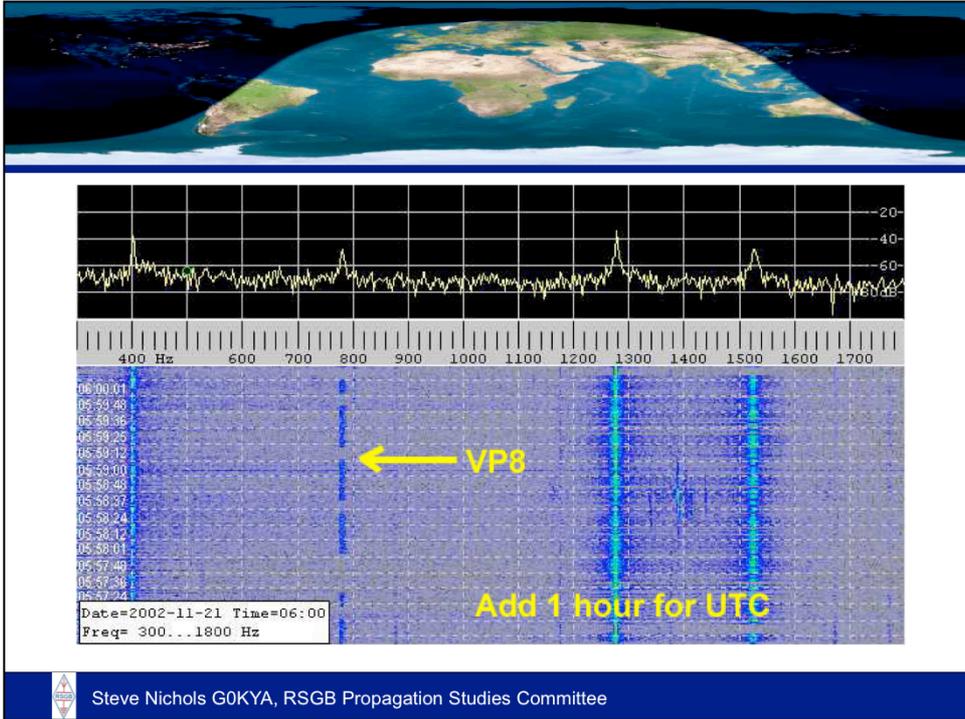
- Mike Harris VP8NO set up a 40m dipole and an Elecraft K2
- Using slow Morse (QRSS) he transmitted his call on 7039.41kHz
- Signals were received in the UK for three weeks using SpectrumLab software

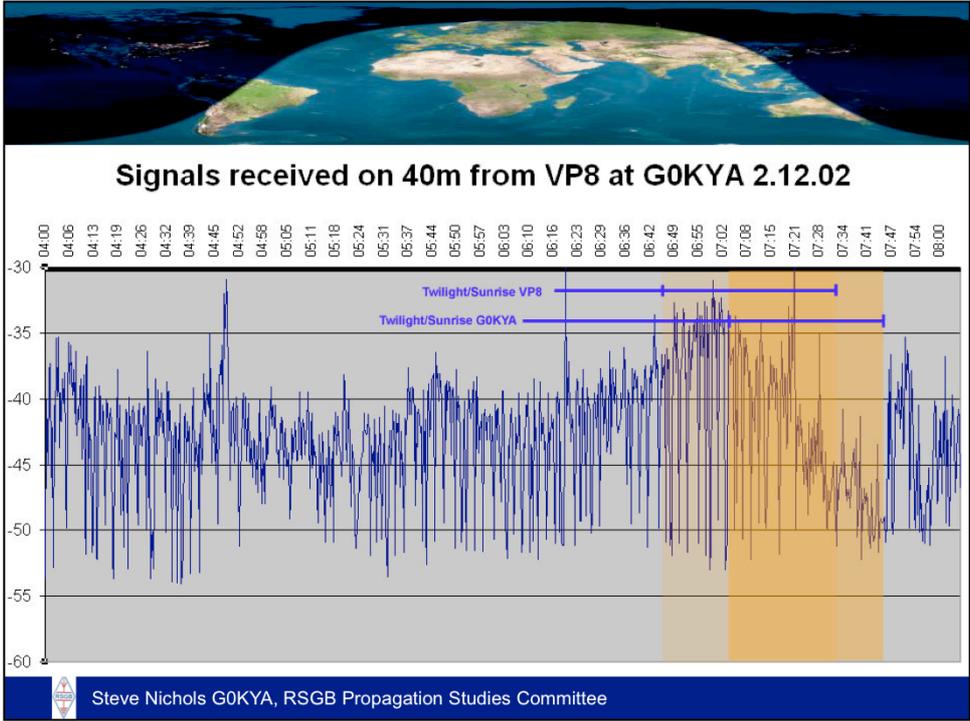


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The G<>VP8 results

- Signals were received through the night
- A greyline enhancement of about 6-10db could be seen about 75% of the time.
- The enhancement occurred in the traditional twilight period - about 30-60 mins. before sunrise.
- This is probably due to a time lag on E/F layer build-up



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The G<>VP8 40m sunset experiment

- Took place throughout May 2003
- Signals were very poor, probably due to the season and high A/K indices.
- There was no confirmed greyline enhancement.



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**But is there an alternative
explanation for contacts
occurring at sunrise/sunset?**



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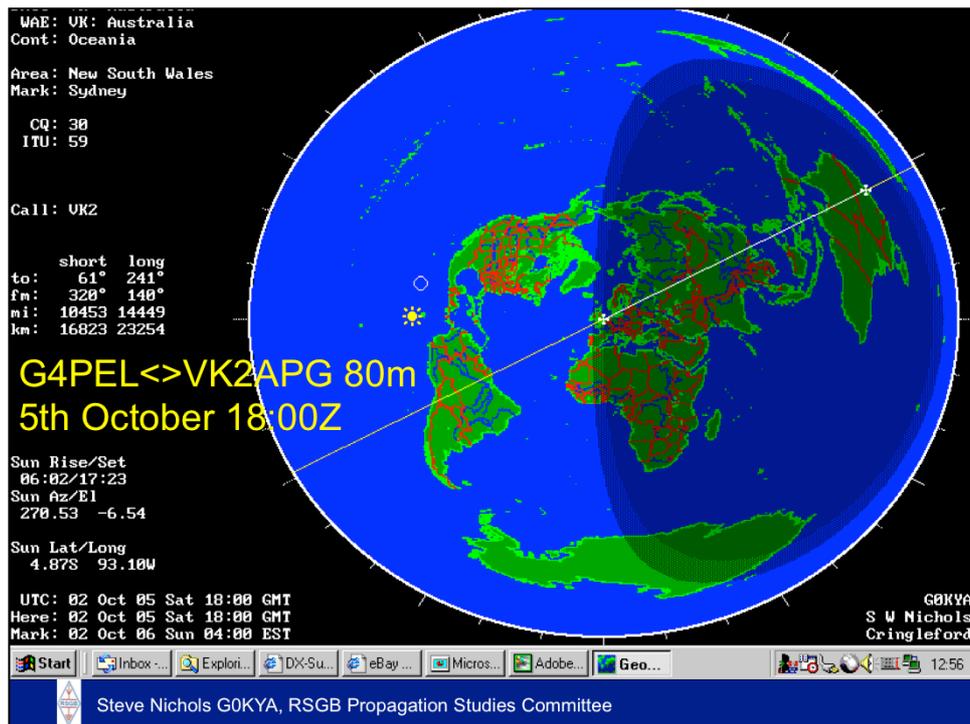


Trans-terminator paths or sunrise/ sunset enhancements

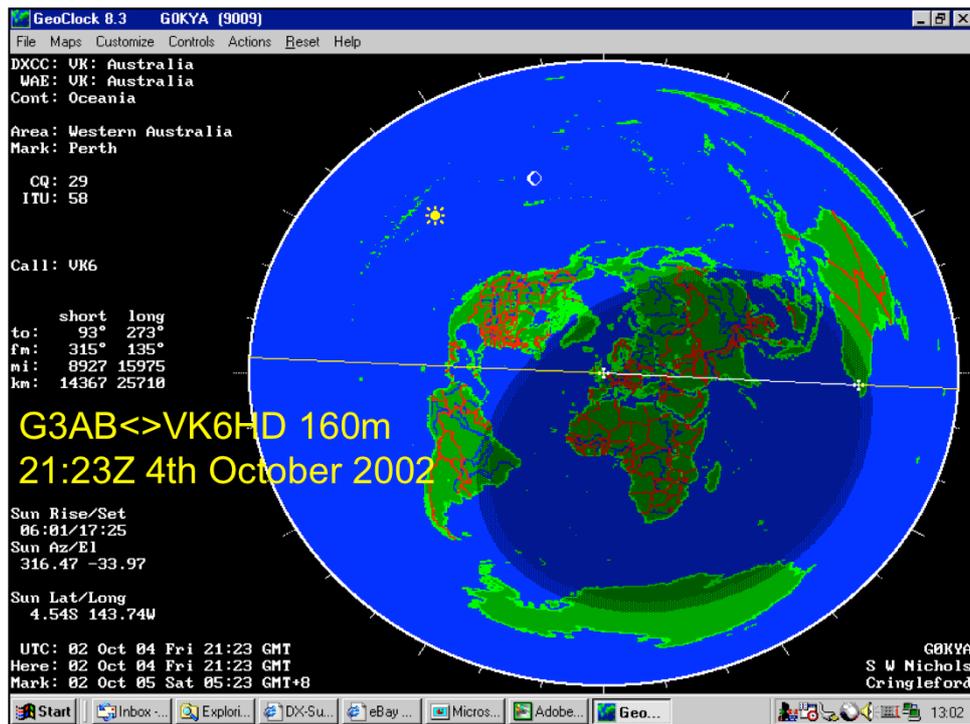
- ON4UN suggests that 160m/80m signals are best at **right angles** to the terminator. Only 40m signals tend to follow a greyline path.
- N4KG says on 80 and 40 meters, signals ALWAYS peak **AFTER SUNRISE** to the west and **BEFORE SUNSET** to the EAST. On 160m, signals can peak at sunrise and sunset and can also peak during darkness. 160m is the **LEAST** predictable band in the amateur spectrum.



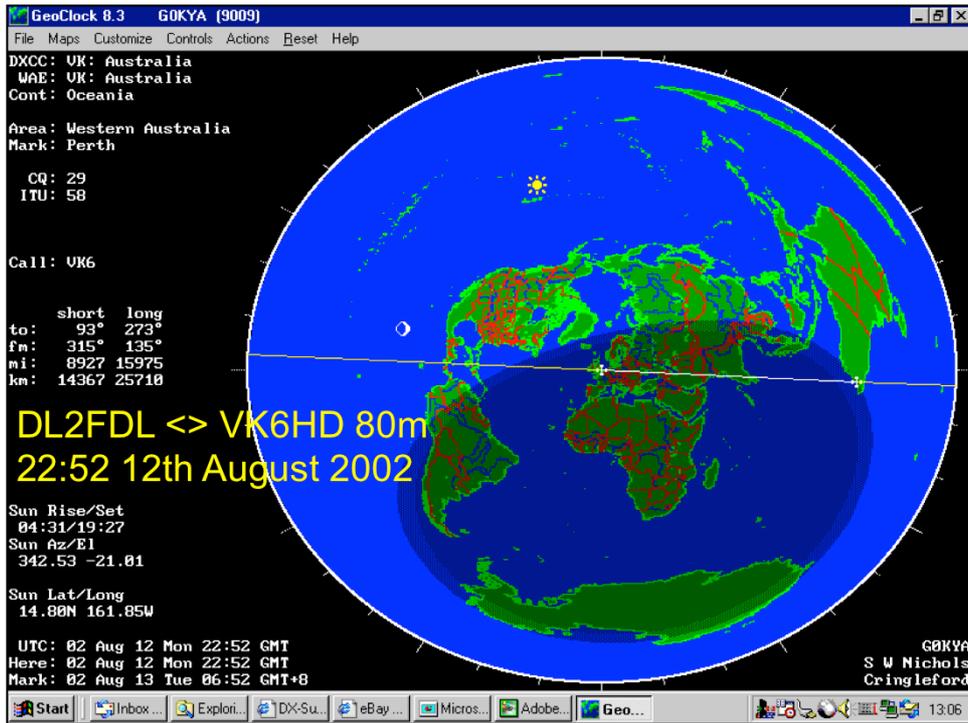
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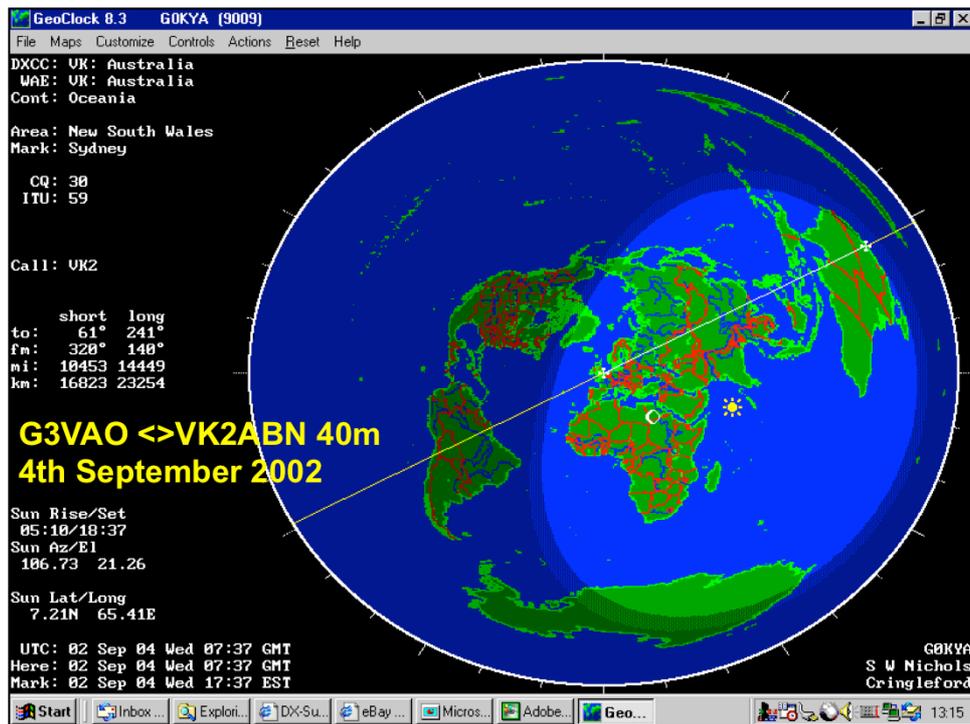


These graphs show that many so-called greyline contacts are anything but. However, they do follow the rule that signals can be strong at local sunrise/sunset from the dark portion of the earth.

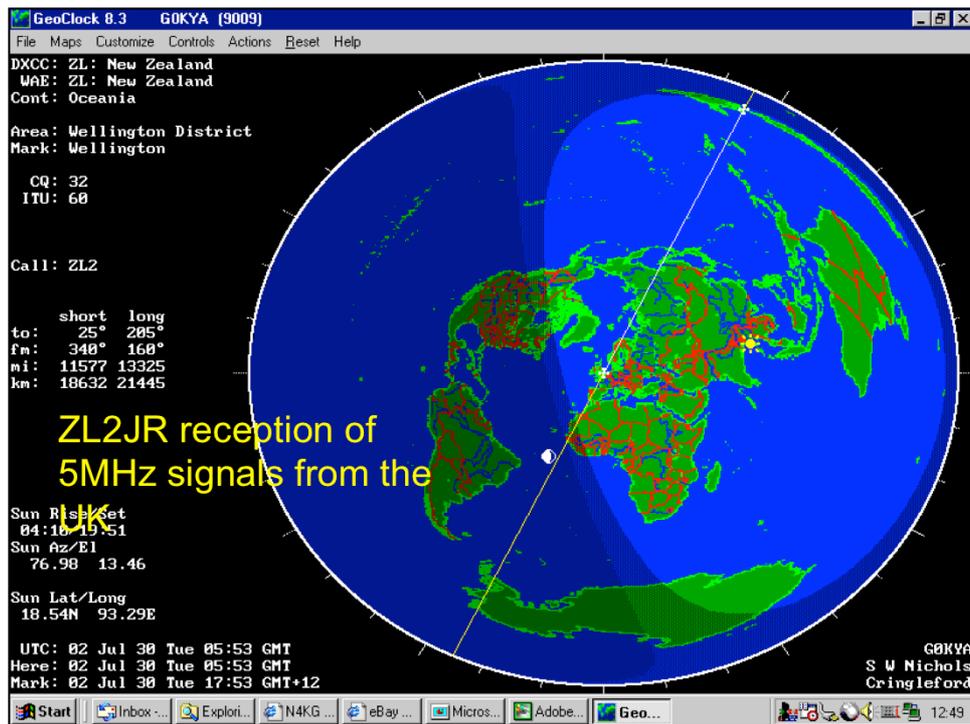


Or at DX sunrise/sunset into the dark portion of the earth.





This is probably due to long path across South America. Note it is a sunset enhancement



This was reported on GB2RS as being greyline. Once again, it is probably longpath. Only by analysing such contacts using GEOCLOCK can you really see what is happening.

GeoClock 8.3 GOKYA (9009)

File Maps Customize Controls Actions Reset Help

DXCC: YB: Indonesia
 WAE: YB: Indonesia
 Cont: Oceania

Area: Indonesia
 Mark: Jakarta

CQ: 28
 ITU: 51 54

Call: YC1

	short	long
to:	82°	262°
fn:	323°	143°
mi:	7198	17713
km:	11578	28586

Sun Rise/Set
 04:33/19:25
 Sun Az/El
 327.54 -17.43

Sun Lat/Long
 14.51N 146.77W

UTC: 02 Aug 13 Tue 21:51 GMT
 Here: 02 Aug 13 Tue 21:51 UTC
 Mark: 02 Aug 14 Wed 04:51 GMT+7

So was this greyline?

GBKYA
 S W Nichols
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start BTopenworl... Outlook ... Windows... Microsoft Po... GeoClock 8... 18:36

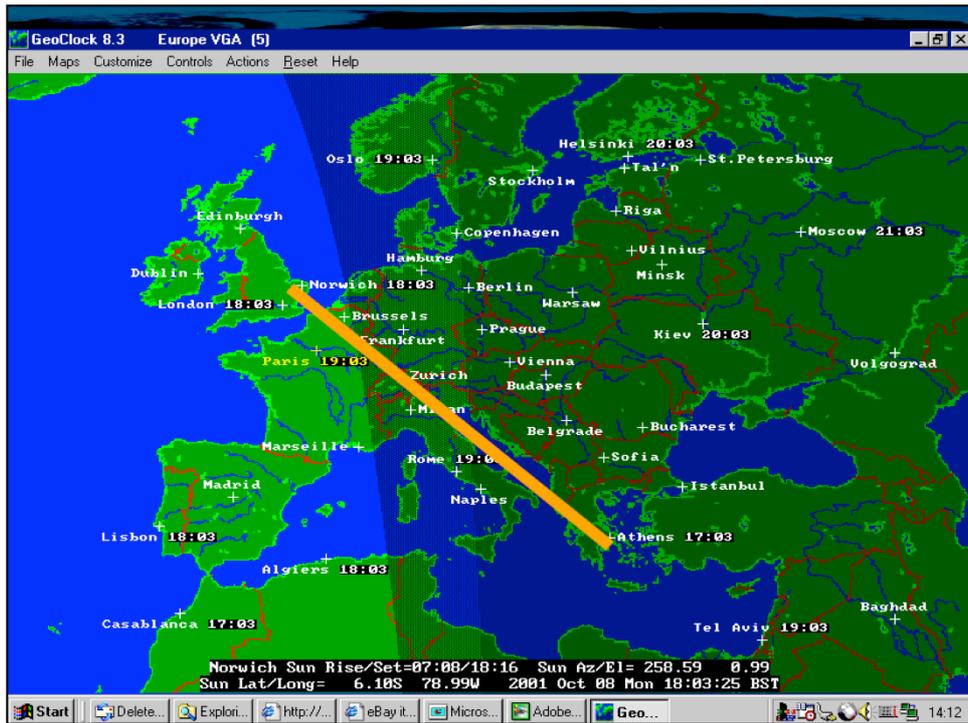


And what about skip focusing on HF - eg. 10m?

This occurs when operating near to the MUF



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The path shown in the previous slide.



In conclusion - true greyline paths

- Greyline paths on 160m/80m are **not** that efficient
- Sunrise/Sunset gives enhancements on the LF bands, but the paths are NOT always “greyline” - they need modelling
- The “corrected” greyline method does not appear to work on all DX paths
- There may be true greyline effects on 40m eg VP8
- D layer absorption due to solar events must be taken into account, as must polar paths



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In conclusion - trans-terminator paths

- On LF, Expect enhancements at your local sunrise and sunset towards the dark zone.
- Also look for enhancements from DX stations across the dark zone at **their** sunset/sunrise, especially perpendicular
- For the above to work you must be in darkness on LF
- Look for skip focusing enhancements on 28Mhz as the terminator crosses the DX path and you are in daylight



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Web addresses for “greyline” information

- **Geoclock** - www.geoclock.com
- **D-layer absorption predictions**
www.sec.noaa.gov/rt_plots/dregion.html
- **Critical frequency (fof2) predictions**
www.spacew.com/www/fof2.html
- **SpectrumLab software** - www.qsl.net/dl4yhf/
- **Beacon Time Wizard** by Taborsoft
www.taborsoft.com
- **RSGB Greyline Research Group** -
<http://uk.groups.yahoo.com/group/greyline/>



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